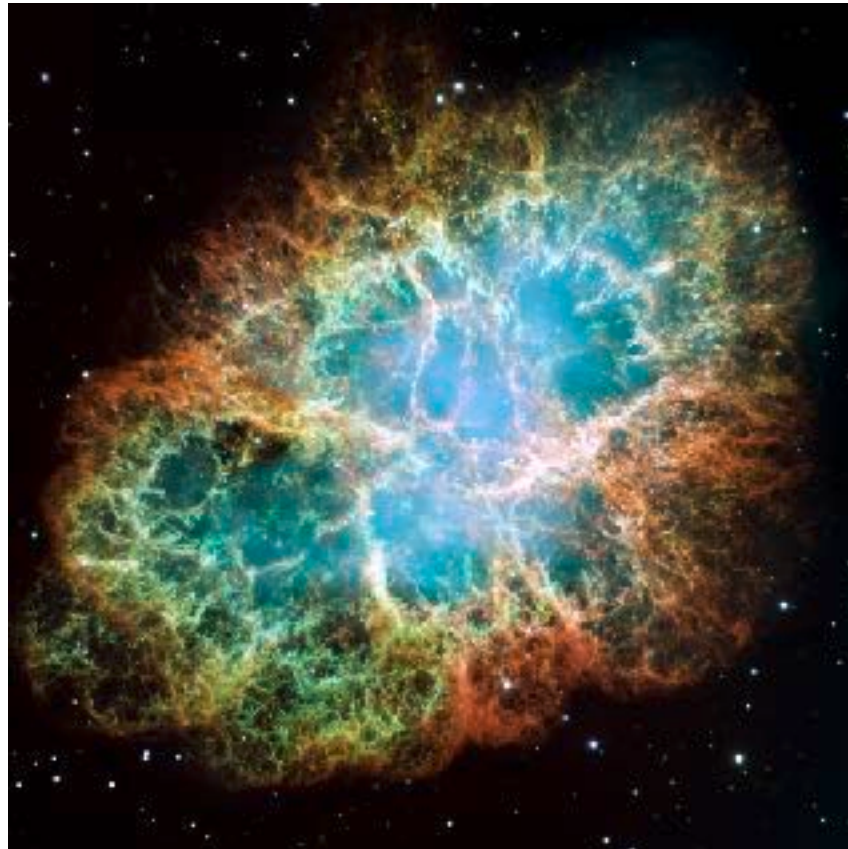


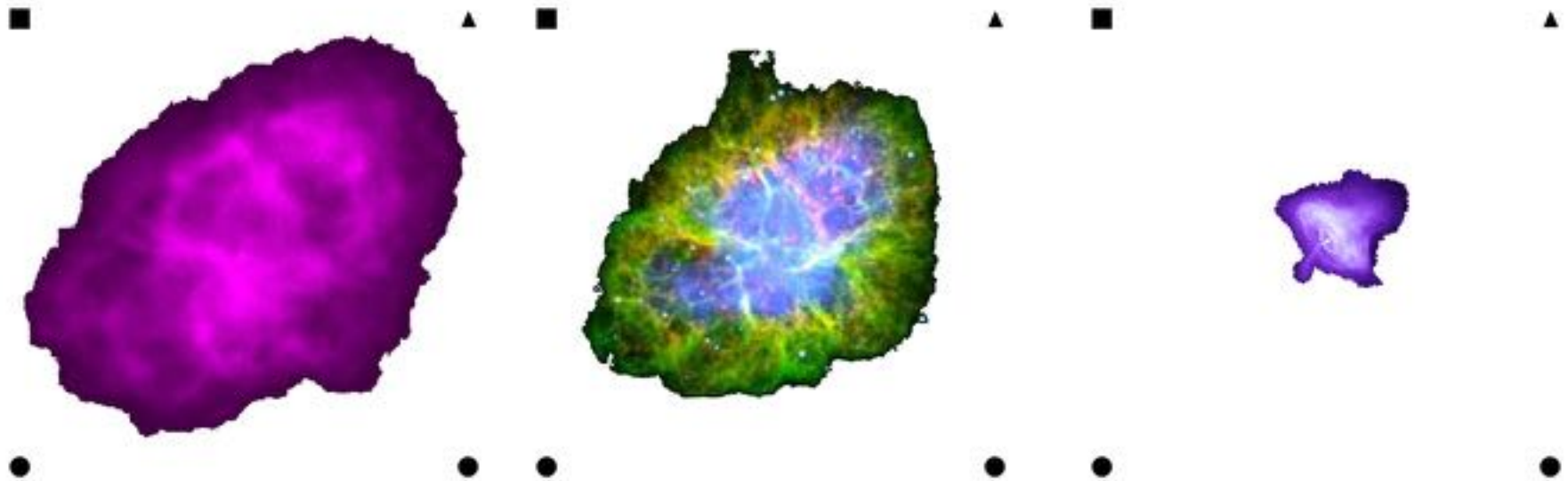
Comments on synchrotron radiation in the Crab



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Multiwavelength observations

Crab synchrotron emission (http://chandra.harvard.edu/edu/formal/composites/crab_overlays.html)



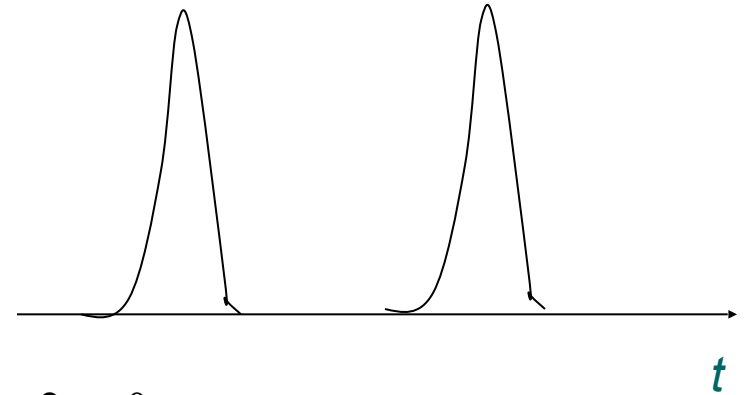
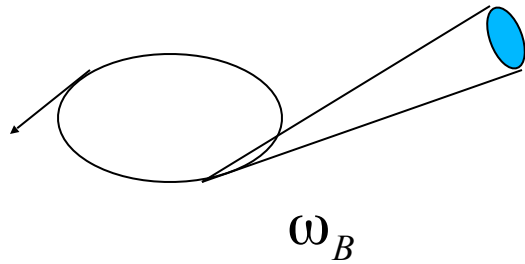
Radio: *extended, low energy electrons*

Optical: *intermediate energy electrons*

X-ray: *compact, close to the source, highest energy electrons*

Synchrotron radiation energies

Emission binned in flashes of light, as conical emission from each electron passes by observer. Characteristic photon frequency scales with square of the Lorentz factor:



$$\omega_c = \frac{3}{2} \gamma^3 \omega_B \sin \alpha = \frac{3}{2} \gamma^2 \omega_B^0 \sin \alpha$$

$$\omega_B \equiv \frac{eB}{\gamma mc} \quad \text{gyration frequency}$$

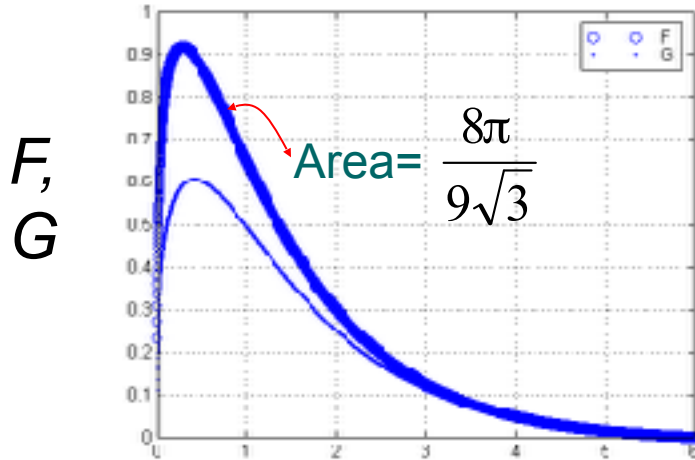
$$\omega_B^0 \equiv \frac{eB}{mc} \quad \text{cyclotron frequency}$$

Synchrotron radiation spectra

Below a critical frequency: spectrum is optically thick (black body)

Above: optically thin (power law from power law energy distribution of electrons)

$$F(x) \sim \frac{8\pi}{9\sqrt{3}} \delta(x-1) \quad \omega_c = \frac{3}{2} \gamma^3 \omega_B \sin \alpha \quad (\text{A})$$

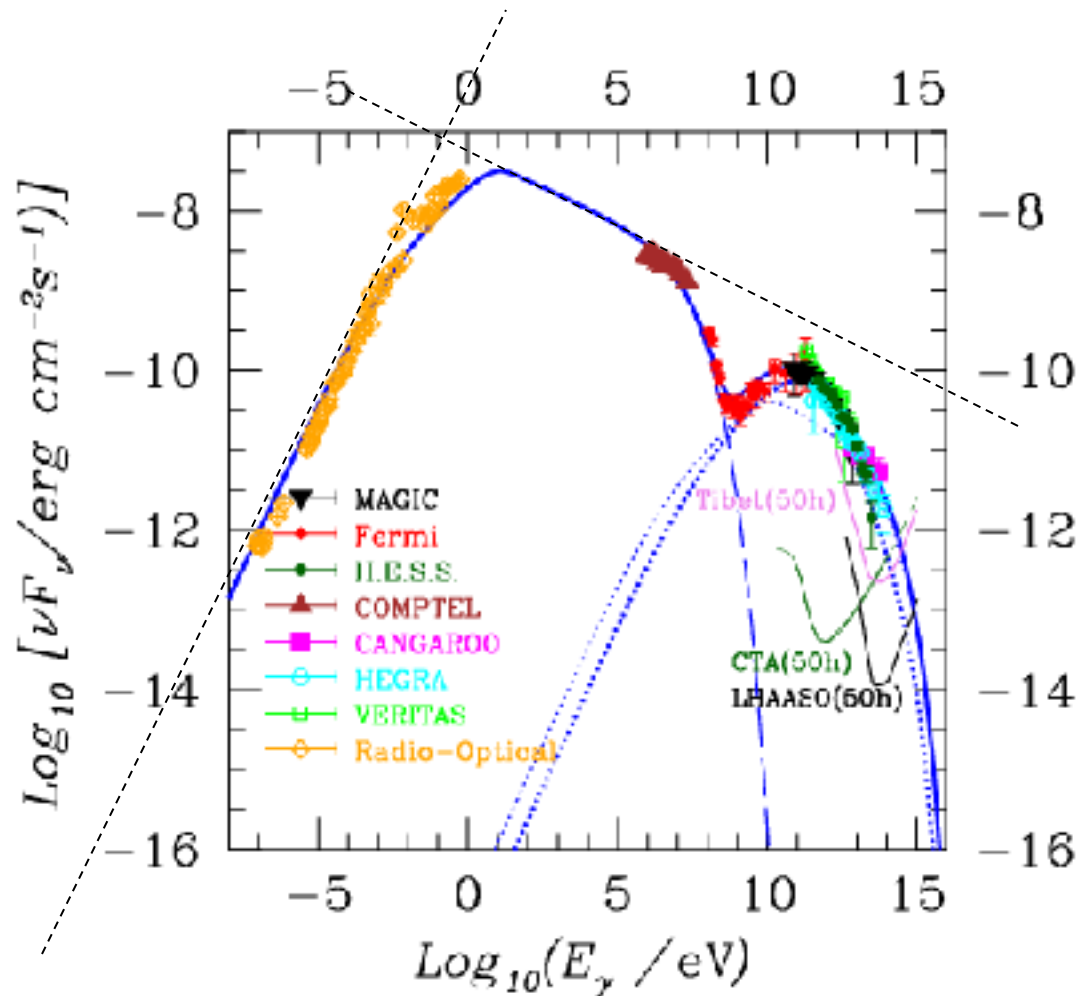


$$\omega_c = 0.29 \times \frac{3}{2} \gamma^3 \omega_B \sin \alpha \quad (\text{B})$$

We commonly use approximation (A)

Synchrotron radiation spectra

Kohri, Yutaka & Kunihiro
 MNRAS, 424, 2249
 Manchester, Staveley-Smith &
 Kesteven, ApJ, 411, 756



*break between radio and
 high energy-energy spectra*

$$\nu_b \approx 10^{14} \text{ Hz}$$

Manchester et al.

$$(\nu_b \approx 2 \times 10^{13} \text{ Hz})$$

----- *extrapolations of radio and high-energy spectra*

Putting it all together:

$$u_B = \frac{1}{8\pi} B^2$$

$$\langle P \rangle_\alpha = \frac{4}{3} \left(\frac{m_e}{m} \right)^2 \sigma_T c \gamma^2 u_B$$

$$t_{syn} = \frac{\gamma m c^2}{\langle P \rangle_\alpha} = \frac{6\pi m_e c^2}{\sigma_T c} \left(\frac{m}{m_e} \right)^3 \frac{1}{\gamma B^2} \cong 10^9 \left(\frac{m}{m_e} \right)^3 \gamma^{-1} B^{-2}$$

$$v_c = \frac{3}{2} \gamma^2 \frac{eB}{mc} : \quad \gamma = \sqrt{\frac{2mc}{3eB} v_c}^{\frac{1}{2}}$$

$$\Rightarrow t_{syn} \cong 5 \times 10^{11} B^{-\frac{3}{2}} v_c^{-\frac{1}{2}} \quad (m = m_e)$$

Results for the Crab:

$$\nu_c \approx 0.2 - 1 \times 10^{14} \text{ Hz} :$$

$$t_{syn} \cong 5 \times 10^{11} B^{-\frac{3}{2}} \nu_c^{-\frac{1}{2}} : \quad B \cong \left(\frac{5 \times 10^{11}}{t_{syn} \nu_c^{\frac{1}{2}}} \right)^{\frac{2}{3}} \cong 0.2 \text{ mG}$$

$$\omega_B^0 = \frac{eBc}{m_e c^2} \cong 3 \text{ kHz}$$

$$\gamma = \sqrt{\frac{2\nu_c}{3\omega_B^0}} \cong 10^5$$